1. We need to satisfy the following two equations:

$$F_{doublet} = F_1 + F_2$$
 $CA_{doublet} = CA_1 + CA_2 = \frac{F_1}{\nu_1} + \frac{F_2}{\nu_2}$

Solution: The refractive efficiencies for ophthalmic crown glass is 58.6 and for dense flint glass is 30.0. From the condition for zero aberration.

$$0 = \frac{F_1}{58.6} + \frac{F_2}{30.0} \tag{1}$$

We also want to satisfy the condition: $F_1 + F_2 = -10$ (2)

 $F_1 = -\frac{58.6 \cdot F_2}{30.0}$ From equation (1):

Sub equation (1) into (2):

$$F_2\left(-\frac{58.6}{30.0}+1\right) = -10 \text{ D} \implies F_2 = 10.49 \text{ D}$$

 $F_1 = -20.49 \text{ D}$ Then it follows that:

The lens might look something like this:



blur size

2. The focal length of an eye with power 60 D for the D-line is: 1.333/60 = 22.22 mm. The radius of curvature of the reduced eve refracting surface (cornea) =

$$F_D = \frac{n_D - 1}{r}$$
 : $r = \frac{n_D - 1}{F_D} = \frac{0.333}{60} = 5.55$ mm

For red light:

$$f_c = \left(\frac{n_C}{n_C - 1}\right)r = 22.32 \text{ mm}$$

For blue light:

$$f_F = \left(\frac{n_F}{n_F - 1}\right)r = 22.01 \text{ mm}$$

Using similar triangles, we can derive the following equation to compute the blur size:

 \Rightarrow blur size = (eye length - focal length) × pupil size $\frac{\text{blur size}}{\text{eye length-focal length}} = \frac{\text{pupil size}}{\text{focal length}}$ focal length

For an 8 mm pupil the C-line and F-line blurs are: 36 and 76.3 microns respectively

For a 2 mm pupil the C-line and F-line blurs are: 9 and 19 microns respectively

You'll notice two things: First, the blue blur is larger then the red blur, which explains the appearance of a bluish halo around white lights. Second, when the pupil is large, the corresponding blurs are also large, which is why these phenomena are especially apparent at night.



5.						
		Exact Approximations		Percentage diff	Percentage differences	
angle in degrees	angle in radians	sin (angle)	1st order	3rd order	1st order	3rd order
0	0.00	0.00	0.00	0.00	0.00	0.00
5	0.09	0.09	0.09	0.09	-0.13	0.00
10	0.17	0.17	0.17	0.17	-0.51	0.00
15	0.26	0.26	0.26	0.26	-1.15	0.00
20	0.35	0.34	0.35	0.34	-2.06	0.01
25	0.44	0.42	0.44	0.42	-3.25	0.03
30	0.52	0.50	0.52	0.50	-4.72	0.07
35	0.61	0.57	0.61	0.57	-6.50	0.12
40	0.70	0.64	0.70	0.64	-8.61	0.21
45	0.79	0.71	0.79	0.70	-11.07	0.35
50	0.87	0.77	0.87	0.76	-13.92	0.54
55	0.96	0.82	0.96	0.81	-17.19	0.81
60	1.05	0.87	1.05	0.86	-20.92	1.18
65	1.13	0.91	1.13	0.89	-25.17	1.68
70	1.22	0.94	1.22	0.92	-30.01	2.33
75	1.31	0.97	1.31	0.94	-35.52	3.18
80	1.40	0.98	1.40	0.94	-41.78	4.29
19.71	0.34	0.34	0.34	0.34	-2.000	
67.65	1.18	0.92	1.18	0.91		2.000







6. For a 20 D lens, the angle at the margins is about 30 degrees. At an angle of about 30 degrees, the Gaussian approximation is off by almost 5 percent. This difference means that the lens makers formula will be close, but not exact. If you form an image with this lens, you will likely observe some blur due to aberrations.

7. Use the radial astigmatism equations here (as long as the final angle is less than 20 degrees you will be OK):

$$P_{t} = P\left[1 + \frac{4\phi^{2}}{3}\right]$$
$$P_{s} = P\left[1 + \frac{\phi^{2}}{3}\right]$$

The equivalent sphere must be -5.5 D, therefore...

$$\frac{P_t + P_s}{2} = -5.5D$$

= $-2.5 \left[1 + \frac{4\phi^2}{3} + 1 + \frac{\phi^2}{3} \right] = -2.5 \left[2 + \frac{5\phi^2}{3} \right] = -5 - \frac{12.5\phi^2}{3}$
 $\Rightarrow 0.5 = \frac{12.5\phi^2}{3}$
 $\Rightarrow \phi = \sqrt{\frac{1.5}{12.5}} = 0.346$ radians = 19.84 degrees

The lens must be tilted 19.84 degrees. Under these conditions, the astigmatism is the difference between the tangential and saggital meridians:

$$P_{t} = -5 \left[1 + \frac{4(0.346)^{2}}{3} \right] = -5.80D$$
$$P_{s} = -5 \left[1 + \frac{0.346^{2}}{3} \right] = -5.20D$$
$$P_{t} - P_{s} = -0.6D$$